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Gauge invariance and superconductivity

## $U(1)$ symmetry breaking and superconducting current

Let's suppose that we multiply all wave functions by the same phase factor  $e^{i\phi}$ . Then field operators transform like  $\psi \rightarrow e^{i\phi}\psi$ . It is easy to see that the normal propagators are invariant with respect to this global  $U(1)$  transformation

$$\langle G \rangle = \langle T\psi\psi \rangle. \text{ On the other hand the anomalous } G_F \text{ is not}$$

$$\langle F \rangle = \langle T\psi\psi \rangle \rightarrow e^{i\phi} \langle T\psi\psi \rangle. \text{ This corresponds to the}$$

physical origin of superconducting phase transition which breaks the  $U(1)$  symmetry. It is less symmetric than normal state. The phase shift requires no energy - therefore it corresponds to

### the Goldstone mode.

What if we consider the local  $U(1)$  symmetry that is assumed that the phase depends on the coordinate? We should take into account the gauge field - vector potential because

the Goldstone mode can be eaten by gauge bosons which becomes massive.

In superconductors it results in Meissner screening or the repulsion of electromagnetic field from superconducting state

To see how that happens let us consider the gauge-invariant one-particle transition probabilities:

### Born-Infeld equations:

$$\left. \begin{aligned} & \left( -i\omega_n + \frac{\hat{p}_x^2}{2m} - \mu \right) G_i + \Delta F^+ = \delta(u_1 - u_2) \\ & \left( -i\omega_n - \frac{\hat{p}_x^2}{2m} - \mu \right) F^+ + \Delta^* G_i = 0 \end{aligned} \right| \quad \text{Here } \hat{p} = -i\tau^L \frac{e}{mc} A - \tau_3 C = \tau_3 \vec{p} \quad V_x = \tau_3 \vec{p} \cdot \vec{A}$$

The gap function  $\Delta$  can have space dependent phase  $\Delta = |\Delta| e^{i\frac{\phi}{2}}$

Let's eliminate it by local gauge transformation

$$G_i(u_1, u_2) = e^{-i\frac{\phi(u_1)}{2} + i\frac{\phi(u_2)}{2}} \tilde{G}_i(u_1, u_2)$$

$$F^+(u_1, u_2) = e^{+i\frac{\phi(u_1)}{2} + i\frac{\phi(u_2)}{2}} \tilde{F}^+(u_1, u_2)$$

Then for the new functions we get

$$\left. \begin{aligned} & \left( -i\omega_n + \frac{(\vec{p}_x + \vec{v}_s)^2}{2m} - \mu \right) \tilde{G}_i + |\Delta| \tilde{F}^+ = \delta(u_1 - u_2) \\ & \left( -i\omega_n - \frac{(\vec{p}_x - \vec{v}_s)^2}{2m} - \mu \right) \tilde{F}^+ + |\Delta| \tilde{G}_i = 0 \end{aligned} \right| \quad \text{Here } \vec{p}_s = -i\vec{\tau}_L$$

$$2mV_s = (\vec{v}_s - \frac{2e}{mc} \vec{A})$$

condensate (Cooper pair)  
velocity

Assuming that  $|A|$  is space-independent we can use momentum repr.

$$G(\mu_1, \mu_2) = G e^{i\vec{p}(\mu_1 - \mu_2)}$$

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If velocity is small, then

$$\frac{(\vec{p} + m\vec{v}_s)^2}{2m} - \mu = \epsilon_p + \vec{p} \cdot \vec{v}_s - \text{the Doppler shift of energy}$$

Hence in Gor'kov Eqs. we can just substitute

$$\omega_n \rightarrow \tilde{\omega}_n + i\vec{p} \cdot \vec{v}_s$$

Current

In order to find how electromagnetic field is modified we need the expression for current.

The current operator in second-q. form

$$\hat{j}(\mu) = \frac{e}{m} \left( \psi^+ (\hat{p} - \frac{e}{c} A) \psi^- - (\hat{p} + \frac{e}{c} A) \psi^+ \psi^- \right)$$

Since we made the gauge transform  $\psi \rightarrow e^{i\frac{q}{2}\mu} \psi$  we get

$$\hat{j}(\mu) = \frac{e}{m} (\psi^+ \hat{p} \psi^- - \hat{p} \psi^+ \psi^-) + 2eV_s \psi^+ \psi^- \rightarrow$$

The average is

$$\overline{\hat{j}(\mu)} = \frac{e}{m} (\overline{\hat{p} + m\vec{v}_s}) G(\mu_1, \mu_2, \tau=0) \Big|_{\mu_1=\mu_2}$$

$$G(\tau=0) = T \sum G_i(\omega_n)$$

careful about  
 $\tau \rightarrow 0$  limit

In momentum representation

$$\overline{\hat{j}} = \frac{e}{m} T \sum \overline{(\hat{p} + m\vec{v}_s)} G(\vec{p}, \omega_n)$$

$$\text{Substitute here } G_i = \frac{\epsilon_p + i\tilde{\omega}_n}{\epsilon_p^2 + \tilde{\omega}_n^2 + i\Delta\epsilon^2} = -\frac{1}{2} \left( \frac{1 + \frac{\epsilon_i}{\epsilon}}{i\tilde{\omega}_n - \epsilon} + \frac{1 - \frac{\epsilon_i}{\epsilon}}{i\tilde{\omega}_n + \epsilon} \right)$$

To calculate the sum over Matsubara frequencies we use

$$T \sum_n \frac{e^{i\omega_n}}{i\omega_n - \epsilon} = f_0(\epsilon) \quad \text{where } f_0 = (e^{\epsilon/k_B T})^{-1}$$

$$\text{Then } j = \frac{e}{m} \sum \overline{(\hat{p} + m\vec{v}_s)} \left[ \left( 1 + \frac{\epsilon_i}{\epsilon} \right) f_0 \left( \frac{\epsilon + p \cdot \vec{v}_s}{T} \right) + \left( 1 - \frac{\epsilon_i}{\epsilon} \right) f_0 \left( \frac{-\epsilon + p \cdot \vec{v}_s}{T} \right) \right] \frac{d^3 p}{(2\pi)^3}$$

For simplicity send  $T \rightarrow 0$  then  $f_0 \rightarrow 0$

$$j = j = eN\vec{v}_s \quad \text{where } N = \int \frac{d^3 k}{(2\pi)^3} = \text{density.}$$

+ Maxwell not  $D = \epsilon \vec{E}$ :

$$\text{If } B \parallel \vec{x}, \quad A \parallel \vec{y} \quad \text{then } \frac{d^2}{d\vec{z}^2} A = -4\pi j = -\frac{4\pi eN}{m} (\nabla \phi - \epsilon A)$$

$b_n^2 = \sqrt{\frac{m}{4\pi eN}}$  - Debye length

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The average is

$$j(u) = -\frac{e}{m} \left[ \hat{p}_1 - \hat{p}_2 + 2mv_s \right] G_1(u_1, u_2) \Big|_{u_1=u_2}$$

In the momentum representation

$$G_1(u_1, u_2) = \int G_1(p_1, p_2) e^{i p_1 u_1 - i p_2 u_2} \frac{d^3 p_1 d^3 p_2}{(2\pi)^6}$$

$$j(k) = -\frac{2e}{m} \int (p + mv_s) G_1(p + k_s, p - k_s) \frac{d^3 p}{(2\pi)^3}$$

In the homogeneous case, the long-wave limit

$$G_1(p + k_s, p - k_s) = G_1(p) \delta(k_s)$$

We assume that the scale on  $v_s$  is much larger than the coherence length of the scale of  $G_1(p)$ . Then  $j(k) \rightarrow$  space-independent

Hence  $j = -\frac{2e}{m} \int (p + mv_s) G_1(p) \frac{d^3 p}{(2\pi)^3} = -\frac{2e}{m} T \int (p + mv_s) G_1(p, \omega_n) \frac{d^3 p}{(2\pi)^3}$

Substituting

$$G_1 = \frac{\xi_p + i\tilde{\omega}_n}{\xi_p^2 + \tilde{\omega}_n^2 + \Delta^2}, \quad \tilde{\omega}_n = \omega_n + i\rho v_s$$

We will often use the relation  $T \leq \frac{e^{i\omega_n}}{i\omega_n - \Delta} = f_0(\Delta/T)$ , where  $f_0 = \frac{1}{e^{\Delta/T} - 1}$ . Fermi dist.

$$\text{If we choose } T \leq \frac{e^{-i\omega_n}}{i\omega_n - \Delta} = f_0(\Delta/T) - 1$$

(a) The second term can be calculated as

$$G_1 = T \leq \frac{e^{i\omega_n}}{i\omega_n^2 + \tilde{\omega}_n^2 + \Delta^2} = T \leq \frac{1}{2} \left( \frac{\xi + \tilde{\omega}/\xi}{i\omega_n - \Delta} + \frac{\xi - \tilde{\omega}/\xi}{i\omega_n + \Delta} \right) = -(\Delta + \tilde{\omega}/\xi) \frac{f_0(\Delta/T)}{2} - (\Delta - \tilde{\omega}/\xi) \frac{f_0'(\Delta/T)}{2}$$

Density of electrons  $\kappa = -2G_1(T=0)$   $\int j_2 = e n_0 v_s$   $\underbrace{\left( \frac{df_0}{d\xi} \cdot \frac{\Delta f_0}{\xi} = 0 \right)}$  odd function

(b) The first term

$$T \leq \int p G_1(p, \omega_n) \frac{d^3 p}{(2\pi)^3} = T \leq \int \frac{\bar{p} (\xi_p + i\tilde{\omega}_n)}{\xi_p^2 + \tilde{\omega}_n^2 + \Delta^2} \frac{d^3 p}{(2\pi)^3}$$

We would like to substitute  $\frac{d^3 p}{(2\pi)^3} \rightarrow \frac{\sin \theta d\theta p^2 dP}{(2\pi)^2} =$

$$= \cos \theta \frac{J_0 d\xi p}{2}$$

$$\text{where } J_0 = \frac{m P_F}{2\pi^2}$$

We can take the integral by  $\epsilon_p$  taking into account the pole which is close to Fermi surface (4)

$$\int \frac{\epsilon_p + i\tilde{\omega}_n}{\epsilon_p^2 + \tilde{\omega}_n^2 + (\Delta l)^2} \frac{d^3 p}{(2\pi)^3} = \frac{\partial_0}{2} \int_{F.S.} \frac{\bar{p}(\epsilon_p + i\tilde{\omega}_n)}{\epsilon_p^2 + \tilde{\omega}_n^2 + (\Delta l)^2} d\epsilon_p d\omega \theta +$$

$$+ \int \frac{\bar{p}}{\epsilon_p - i\tilde{\omega}_n} \frac{d^3 p}{(2\pi)^3}$$

the Fermi sea, when

Off-shelf contribution

Main contribution from  $p$  far from Fermi surface.

$$\int_0^\infty \int_0^{2\pi} p^3 dp \left( \sin \theta d\theta \right) \frac{\cos \theta}{\epsilon_p - i\omega_n + p \sin \theta} = - \int_0^\infty p^3 dp \int_0^{2\pi} \frac{p V_s \omega \sin \theta d\theta}{(\epsilon_p - i\omega_n)^2}$$

$$= -\frac{2}{3} \sqrt{s} \int_0^\infty \frac{p^4 dp}{(\epsilon_p - i\omega_n)^2}$$

$$\frac{T}{(2\pi)^2} \leq -\frac{2}{3} \sqrt{s} \int_0^\infty \frac{p^4 dp}{(\epsilon_p - i\omega_n)^2} = -\frac{2}{3} \frac{\sqrt{s}}{(2\pi)^2} \int_0^\infty p^4 \frac{d}{dp} f_0\left(\frac{\epsilon_p}{T}\right) dp = -\frac{2m}{3} \sqrt{s} \frac{p_F^3}{(2\pi)^2} = -\frac{nm\sqrt{s}}{2}$$

$$\text{since } n = \frac{p_F^3}{3\pi^2}$$

Hence finally

$$T \leq \left( \bar{p} G_1(p, \omega_n) \frac{d^3 p}{(2\pi)^3} \right) = i\pi \int_0^\infty \frac{\bar{p}_F \tilde{\omega}_n T}{(\tilde{\omega}_n^2 + (\Delta l)^2)^{1/2}} d\omega \theta = \epsilon n \sqrt{s}$$

The current

$$j = -\frac{ie}{m} \bar{n} \int_0^\infty \frac{\bar{p}_F \tilde{\omega}_n T}{(\tilde{\omega}_n^2 + (\Delta l)^2)^{1/2}} d\omega \theta \quad \tilde{\omega}_n = \omega_n + i\bar{p}_F \sqrt{s}$$

Expand up to the first order in  $\sqrt{s}$

$$\frac{\bar{p}_F \tilde{\omega}_n}{(\tilde{\omega}_n^2 + (\Delta l)^2)^{1/2}} = \frac{\bar{p}_F \omega_n}{(\omega_n^2 + (\Delta l)^2)^{1/2}} + \frac{\bar{p}_F i(\bar{p}_F \sqrt{s})}{(\omega_n^2 + (\Delta l)^2)^{1/2}} - \frac{i \bar{p}_F \omega_n^2 (\bar{p}_F \sqrt{s})}{(\omega_n^2 + (\Delta l)^2)^{3/2}} \Rightarrow$$

$$\Rightarrow \frac{\bar{p}_F (\bar{p}_F \sqrt{s}) (\Delta l)^2}{(\omega_n^2 + (\Delta l)^2)^{3/2}}$$

$$j = t \frac{2e}{m} \bar{n} \int_0^\infty \sqrt{s} p_F^2 2 \int_0^1 x^2 dx \lesssim \frac{T |\Delta l|^2}{(\omega_n^2 + (\Delta l)^2)^{3/2}} = \frac{t q}{3} \frac{e}{m} \bar{n} 2 \bar{p}_F^2 \lesssim \frac{T |\Delta l|^2}{(\omega_n^2 + (\Delta l)^2)^{3/2}} \sqrt{s}$$

Recall that  $\mathcal{D}_0 = \frac{mP}{2\pi k}$  then  $2\frac{\mathcal{D}_0 P^2}{z m} = n_0$  (5)

So that

$$j = ze n_0 \leq \frac{T |A|^2 \pi}{(w_n^2 + |A|^2)^{3/2}} \quad \sqrt{s} = + \frac{ze n_0}{2m} \left( T \varphi - \frac{ze}{mc} A \right) \leq \frac{T |A|^2 \pi}{(w_n^2 + |A|^2)^{3/2}}$$

$$= \frac{ze^2 n_0}{mc^2} \left( \frac{ze}{2e} \nabla \varphi - A \right) \leq \frac{T |A|^2 \pi}{(w_n^2 + |A|^2)^{3/2}} = \frac{e}{4\pi} b_n^{-2} \left( \frac{bc}{ze} \nabla \varphi - A \right)$$

where  $b_n^{-2} = \frac{4\pi n_s e^2}{mc^2}$

$n_s = n_0 |A|^2 \leq \frac{\pi}{(w_n^2 + |A|^2)^{3/2}}$  - density of superconducting component

so that  $j = e n_s \sqrt{s}$  - definition of  $n_s$

London model  $T \varphi = 0$  then  $j = - \frac{c}{4\pi} b_n^{-2} A$

Maxwell eq.  $\nabla \cdot B = - \frac{c}{4\pi} A$

$\nabla \cdot \nabla A = - \frac{c}{4\pi} A$

Let's consider 1-d problem

$$\vec{B}_0 \uparrow \quad \text{Super.} \quad \vec{A} = A_y(x) \hat{i} \quad \text{not } \nabla A = - \vec{y} \frac{d^2}{dx^2} A_y$$

$$\text{vacuum} \quad B_z = \frac{\partial A_y}{\partial x} \quad \frac{d^2}{dx^2} A_y - \frac{c}{4\pi} A_y = 0$$

$$A_y = A_0 e^{-x/l_n}$$

$$B_z = B_0 e^{-x/l_n}$$

gauge field  
finite mass

Meissner effect

Physically that means that when we apply magnetic field it induces current along the surface. The field created by this current exactly compensates external field

# Density of superconducting component

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$$h_s = \mu_0 |\Delta|^2 \leq \frac{\pi T}{n} \frac{1}{(\omega_n^2 + |\Delta|^2)^{1/2}}$$

Let's consider two limiting cases

a)  $T \rightarrow T_c$        $h_s = \mu_0 |\Delta|^2 \leq \frac{\pi T}{\omega_n^3} = \mu_0 |\Delta|^2 \leq \frac{1}{(2n+1)^3 (\pi)^2} =$   
 $|\Delta| < \omega_n$

$$= \mu_0 \frac{|\Delta|^2}{T_c^2} \quad \frac{\cancel{\pi} \cancel{(2n+1)^3}}{\cancel{8} T_c^2} \quad \sim (1 - \frac{T}{T_c}) \quad \Delta \rightarrow \infty$$

b)  $T \rightarrow 0$        $\frac{h_s}{\mu_0} = |\Delta|^2 \int_{-\infty}^{\infty} \frac{\pi T \, dn}{((2n+1)^2 (\pi T)^2 + |\Delta|^2)^{1/2}} = \int_0^{\infty} \frac{dx}{x^2 + b^2} = 1$

$$x = \sinh t \\ dx = \cosh dt$$

$$\int \frac{dt}{\cosh^2 t} = \tanh t$$

$$(x^2 + b^2)^{1/2} = \sqrt{b^2 t}$$

Flux quantization.

(2)

London approximation  $\nabla \phi = 0$  is not always correct

Phase gradient can compensate the vector potential

Consider multiply connected domain.

Due to the Meissner effect there is no current in the bulk  $j_s = 0$



$$\frac{ie}{2e} \nabla \phi - A = 0$$

$$S \nabla \phi = 2\pi n \Rightarrow \oint A dl = \Phi = \frac{ie}{2e} \cdot 2\pi n = \frac{\pi hc}{e}$$

$$\Phi_0 = \frac{\pi hc}{e} = \text{magnetic flux quantum}$$

Magnetic field is confined within non-superconducting domain. Such domains can be created either artificially just drilling holes in superconductor. Also they can appear naturally when the non-superconducting region is just the normal metal. Such normal tubes with localized magnetic flux are Abrikosov vortices